

Neutron-antineutron oscillation improvements and baryogenesis

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Apologies for few references. Many in our paper – please consult

Outline

Introduction

NNbar oscillations for free neutrons

EFT for Nnbar oscillations

Minimal EFT for $\Delta B=2$ baryogenesis

Calculations of baryon asymmetry and n-nbar oscillations

Conclusion

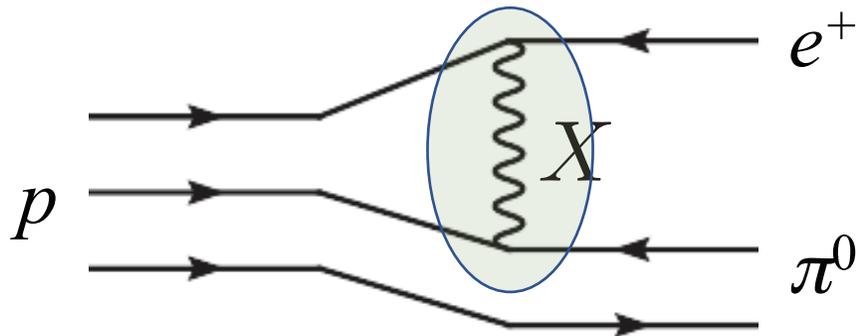
B and L are accidental symmetries of SM – subject to violation

Proton decay can occur by higher-dim operator.

$$\frac{1}{\Lambda_p^2} (\bar{d}^c \bar{u}^c q \ell) + \dots$$

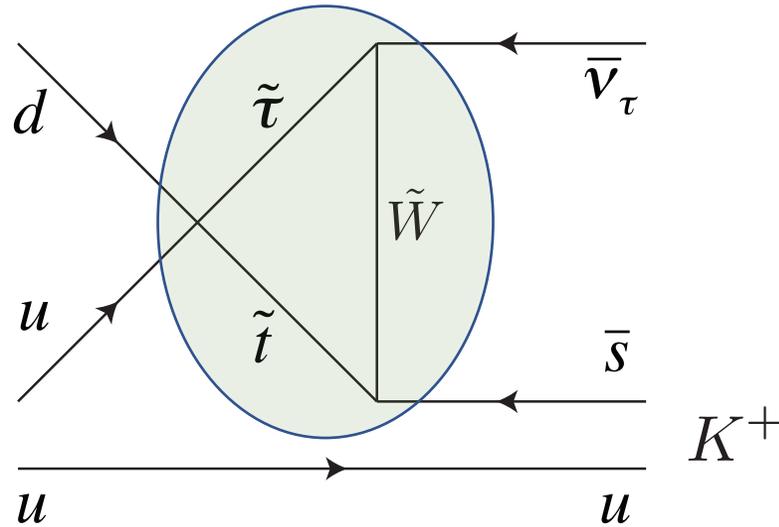
This operator likely exists with Planck suppressed couplings at least. $\Delta B = \Delta L = 1$

This operator could exist with smaller scale suppression in GUT theories, etc..



$$p \rightarrow e^+ \pi^0$$

(SUSY and Non-SUSY GUTs)

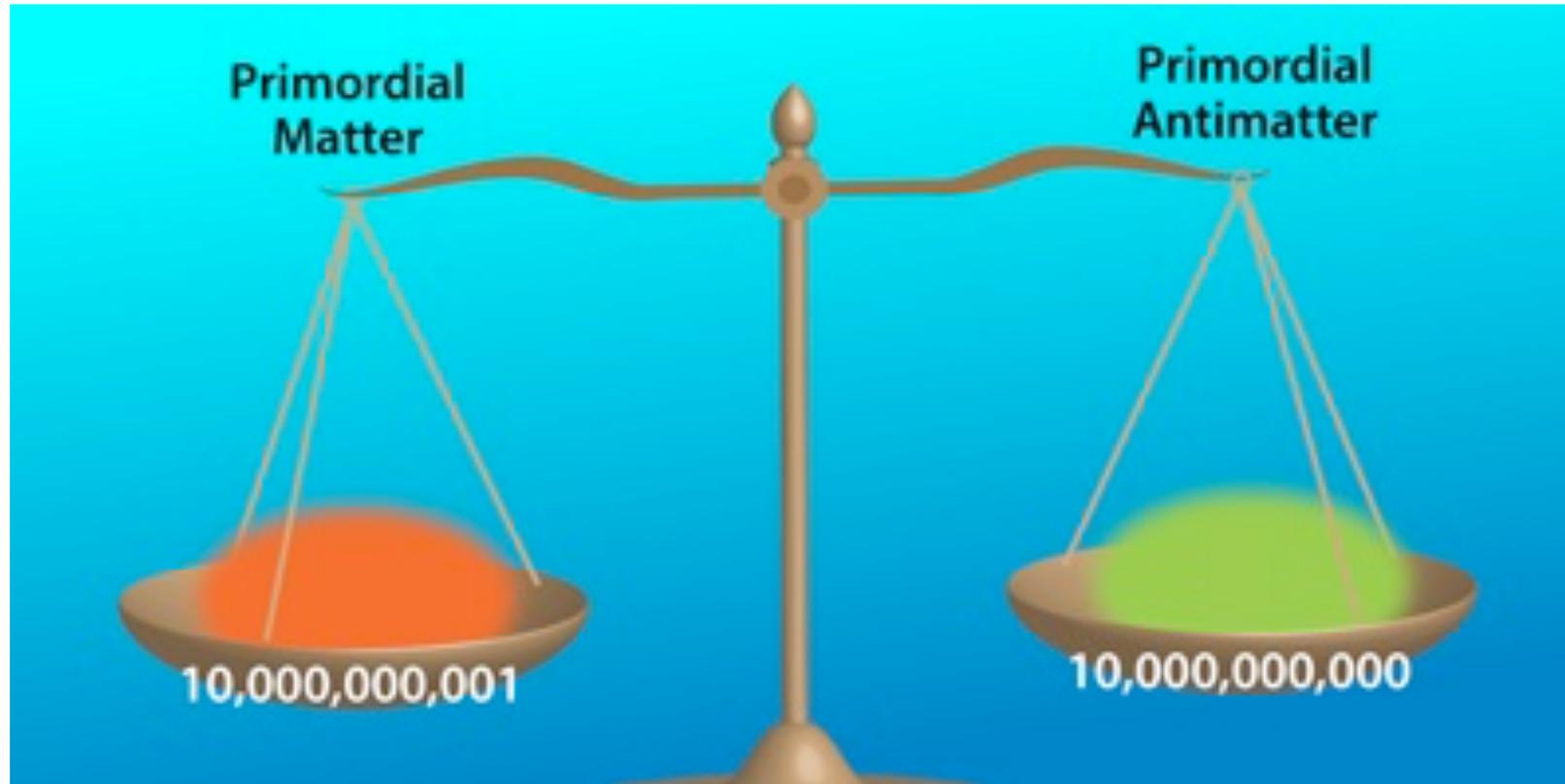


$$p \rightarrow \bar{\nu} K^+$$

(SUSY GUTs)

B, L violation connected to baryon asymmetry puzzle

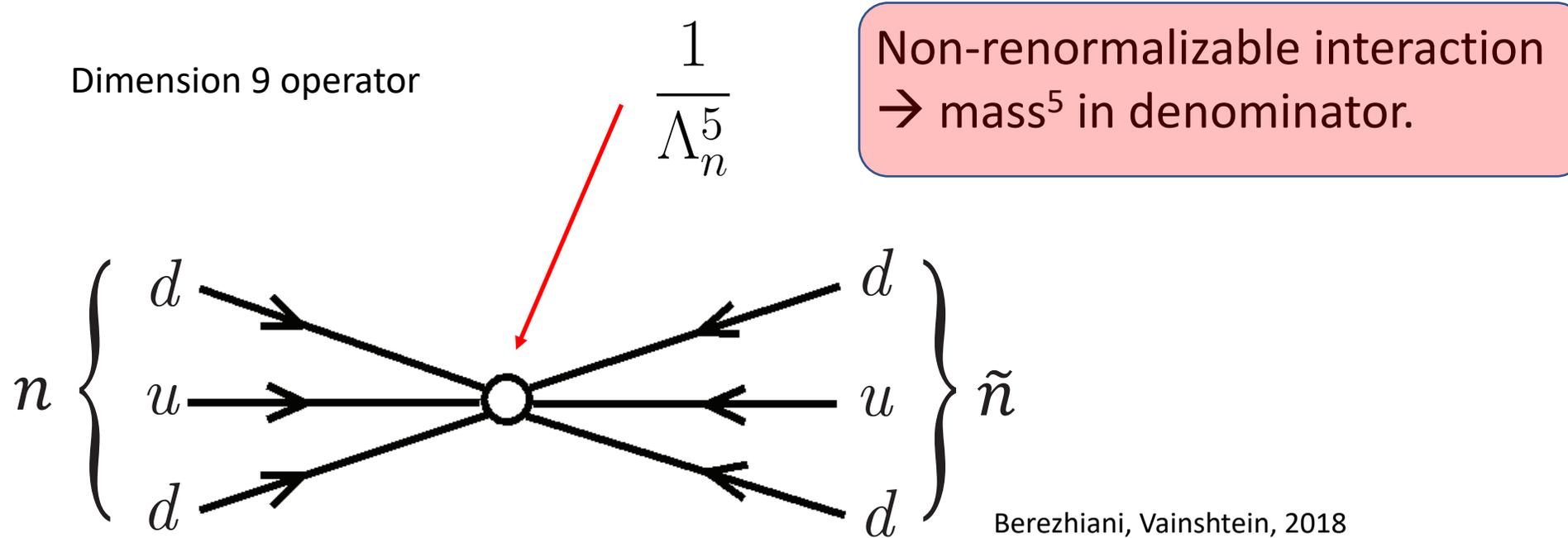
B-L violation required to avoid sphaleron washout above EW scale – proton decay conserves B-L



M.G. Strauss

Let's look to more direct connections to B violation.

$\Delta B=2$ baryon number violating interaction in EFT



Other $|\Delta B| = 2$, $\Delta L = 0$ processes include dinucleon decays: $nn \rightarrow \pi^0\pi^0$, $pp \rightarrow \pi^+\pi^+$, $pn \rightarrow \pi^+\pi^0$ probe the same operators as n - \tilde{n} oscillation, while $pp \rightarrow K^+K^+$ can be relevant for B -violating new physics with suppressed couplings to first-generation quarks.

Past and future experimental sensitivities to oscillation lifetime

- Current free neutron bound $\tau \sim 10^8$ s from ILL, and somewhat better at Super-K
- Future prospects at ESS (free neutrons), DUNE and Hyper-K up to $\tau \sim 10^{9-10}$ s
- Such oscillation times probe new physics scales up to $\Lambda_{\text{new}} \sim (\tau \Lambda_{\text{QCD}}^6)^{1/5} \sim 10^{5-6}$ GeV.
- Probes beyond LHC scales (albeit in only just this way)

Quantum mechanics of neutron-antineutron oscillations

Evolution governed by Schrödinger equation:

$$\mathcal{H}_{\text{eff}}|\psi\rangle = i\frac{\partial}{\partial t}|\psi\rangle$$

Where Hamiltonian given by

$$\begin{aligned}\mathcal{H}_{\text{eff}}|n\rangle &= \left(m_n - i\frac{\Gamma}{2} + \mathcal{E}_n\right)|n\rangle + \delta|\bar{n}\rangle \\ \mathcal{H}_{\text{eff}}|\bar{n}\rangle &= \left(m_n - i\frac{\Gamma}{2} + \mathcal{E}_{\bar{n}}\right)|\bar{n}\rangle + \delta|n\rangle\end{aligned}$$

The matrix $\langle \mathcal{H}_{\text{eff}} \rangle$ in the $\{n, \bar{n}\}$ basis is

$$\langle \mathcal{H}_{\text{eff}} \rangle = \begin{pmatrix} m_n - i\frac{\Gamma}{2} + \mathcal{E}_n & \delta \\ \delta & m_n - i\frac{\Gamma}{2} + \mathcal{E}_{\bar{n}} \end{pmatrix}$$

where m_n is mass of the neutron, Γ is the decay width (i.e., neutron lifetime is $\tau_n = 1/\Gamma$), δ is contribution from \mathcal{H}_{eff} that enables $n \leftrightarrow \bar{n}$ transitions, and \mathcal{E}_n and $\mathcal{E}_{\bar{n}}$ are any other additional contributions to the energy of the n and \bar{n} states respectively. If the neutrons were propagating completely freely in space with no other matter around and no magnetic field, etc., $\mathcal{E}_{n,\bar{n}} = 0$. But since that is never the case in experimental configurations, we must keep this term.

If the neutron and antineutron mix then energy eigenstates (or, “mass eigenstates”) \mathcal{H}_{eff} are mixtures of n and \bar{n} which we denote as n_1 and n_2 :

$$\begin{pmatrix} |n_1\rangle \\ |n_2\rangle \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} |n\rangle \\ |\bar{n}\rangle \end{pmatrix}$$

Assuming $|n_i\rangle$ are eigenstates of \mathcal{H}_{eff}

$$\mathcal{H}_{\text{eff}} |n_i\rangle = E_i |n_i\rangle$$

one then expands evolves the wave function

$$|\psi\rangle(t) = c_1 |n_1\rangle e^{-iE_1 t} + c_2 |n_2\rangle e^{-iE_2 t}$$

Subjecting it to boundary condition $|\psi\rangle(0) = |n\rangle$ we find

$$|\psi\rangle(t) = (\cos^2\theta e^{-iE_1 t} + \sin^2\theta e^{-iE_2 t}) |n\rangle + \cos\theta \sin\theta (e^{-iE_1 t} - e^{-iE_2 t}) |\bar{n}\rangle$$

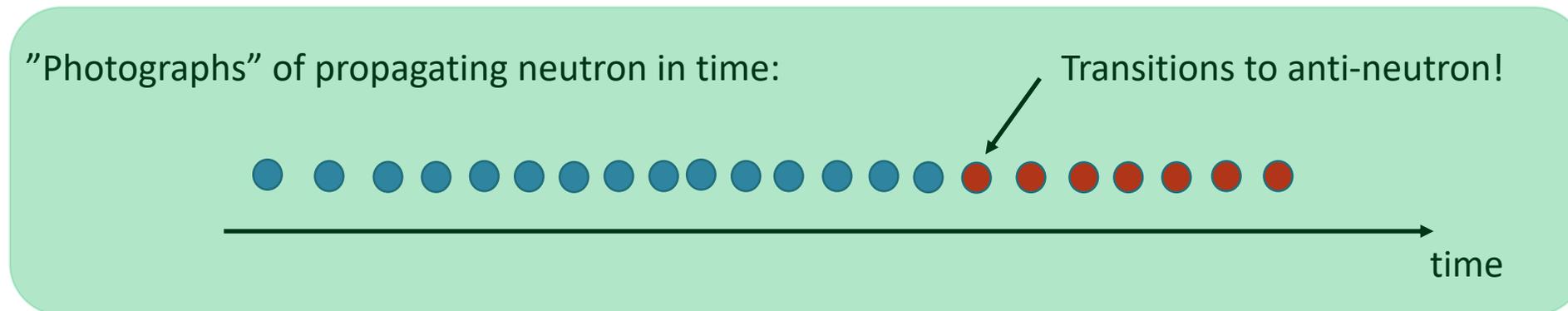
$$|\psi\rangle(t) = (\cos^2 \theta e^{-iE_1 t} + \sin^2 \theta e^{-iE_2 t}) |n\rangle + \cos \theta \sin \theta (e^{-iE_1 t} - e^{-iE_2 t}) |\bar{n}\rangle$$

We compute the probability that $|\psi\rangle(t)$ is measured to be a \bar{n} by the standard probability computation in quantum mechanics,

$$P[\bar{n}(t)] = |\langle \bar{n} | \psi \rangle(t)|^2 = e^{-\Gamma t} \sin^2(2\theta) \sin^2\left(\frac{\Delta E t}{2}\right), \quad \text{where,}$$

$$\Gamma = \text{Im}(E_1 + E_2), \quad \text{and} \quad \Delta E = E_1 - E_2.$$

The first term, $e^{-\Gamma t}$, is associated with the lifetime of the neutron.



Approximations valid for reactor environment

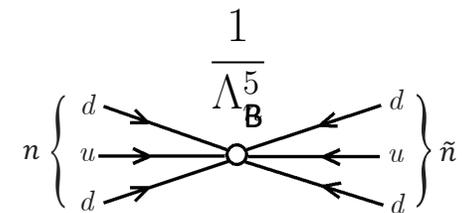
For $m_n \gg |\mathcal{E}_n - \mathcal{E}_{\bar{n}}| \gg \delta$, which will be justified later in the nuclear reactor experimental context, one can make the approximations

$$E_1 \simeq m_n + \mathcal{E}_n - i\frac{\Gamma}{2}, \quad E_2 \simeq m_n + \mathcal{E}_{\bar{n}} - i\frac{\Gamma}{2}, \quad \text{where} \quad \mathcal{E}_n = -\mathcal{E}_{\bar{n}} = -\mu_n \cdot B$$

$$\Delta E = E_1 - E_2 = \mathcal{E}_n - \mathcal{E}_{\bar{n}}, \quad \text{and} \quad \sin 2\theta = \frac{2\delta}{\mathcal{E}_n - \mathcal{E}_{\bar{n}}}.$$

Under these assumptions we can now rewrite the transition probability as

$$P[\bar{n}(t)] = e^{-\Gamma t} \left(\frac{2\delta}{\mathcal{E}_n - \mathcal{E}_{\bar{n}}} \right)^2 \sin^2 \left(\frac{(\mathcal{E}_n - \mathcal{E}_{\bar{n}})t}{2} \right).$$



Recall, δ is the interaction that allows neutron to antineutron transition!

$$\delta = m_{n/11}^6 / \Lambda_B^5$$

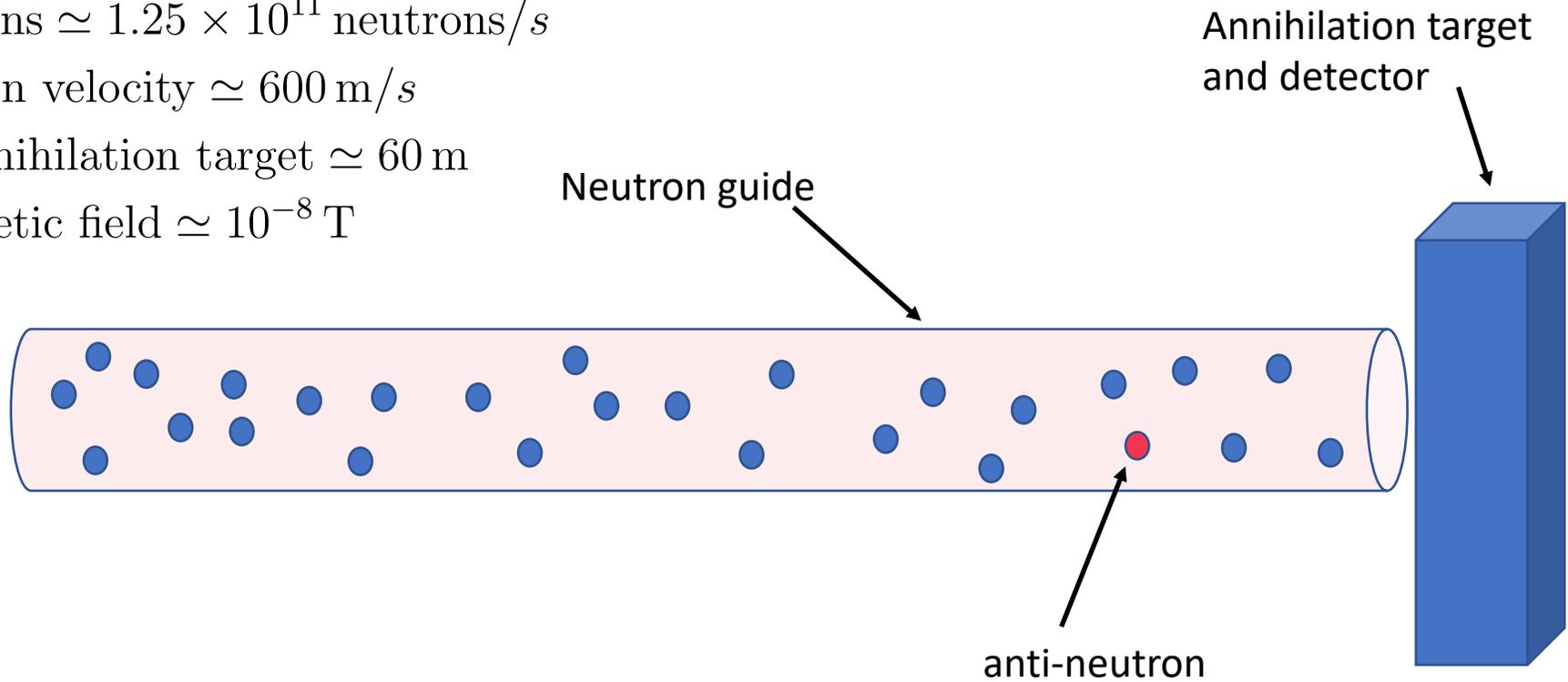
F = Flux of neutrons $\simeq 1.25 \times 10^{11}$ neutrons/s

v_{avg} = average neutron velocity $\simeq 600$ m/s

L = distance to annihilation target $\simeq 60$ m

B = ambient magnetic field $\simeq 10^{-8}$ T

Neutrons
from reactor



From the average velocity data, the average time for the neutron to make it to the annihilation target is $t_{\text{avg}} = L/v_{\text{avg}} \simeq 0.1$ s. This is where the state $|\psi\rangle(t)$ is measured and its wave function collapses to n or \bar{n} , at time $= t_{\text{avg}}$ when it interacts with the annihilation target.

$$P[\bar{n}(t_{\text{avg}})] \simeq \delta^2 t_{\text{avg}}^2 = 10^{-18} \left(\frac{10^8 \text{ s}}{\tau_{n\bar{n}}} \right)^2 \left(\frac{t_{\text{avg}}}{0.1 \text{ s}} \right)^2, \quad \longrightarrow \quad \tau_{n\bar{n}} \simeq (2 \times 10^8 \text{ s}) \left(\frac{F}{1.25 \times 10^{11} \text{ neutrons/s}} \right)^{1/2} \left(\frac{T_{\text{run}}}{1 \text{ yr}} \right)^{1/2}.$$

where $\tau_{n\bar{n}} \equiv 1/\delta$ (oscillation time)

~ Current limit!

Let's explore connection of n - \bar{n} oscillations with baryogenesis

Assume: simple minimal EFT with minimal new particle content that achieves baryogenesis through B violating decays.

These new particles can simultaneously allow n - \bar{n} oscillations \rightarrow correlated

One direction (“Majorana fermion baryogenesis”)

Other directions: EW baryogenesis, Affleck–Dine baryogenesis, leptogenesis, ...

Lowest dimensional operator contributing to n - \bar{n} oscillations is dimension 9

$$\mathcal{O}_{n\bar{n}} \sim (uudddd)$$

We will focus on just one of these operators for illustration, and because it matches the low-scale EFT of the minimal scenario for baryogenesis.

$$\mathcal{L} \supset c_1 \frac{1}{2} \epsilon_{ijk} \epsilon_{i'j'k'} (\bar{u}_i^c P_R d_j) (\bar{u}_{i'}^c P_R d_{j'}) (\bar{d}_k^c P_R d_{k'}) + \text{h.c.},$$

with $c_1 \equiv \left(\Lambda_{n\bar{n}}^{(1)}\right)^{-5}$.

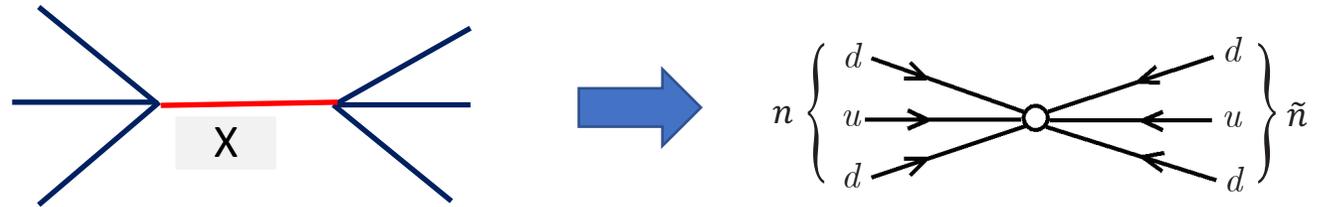
We performed state-of-the-art RG evolution of the operator coefficient

$$\begin{aligned} \frac{c_1(\mu_0)}{c_1(M)} &= \left[\frac{\alpha_s^{(4)}(m_b)}{\alpha_s^{(4)}(\mu_0)} \right]^{\frac{6}{25}} \left[\frac{\alpha_s^{(5)}(m_t)}{\alpha_s^{(5)}(m_b)} \right]^{\frac{6}{23}} \left[\frac{\alpha_s^{(6)}(M)}{\alpha_s^{(6)}(m_t)} \right]^{\frac{2}{7}} \\ &= \{0.726, 0.684, 0.651, 0.624\}, \\ &\text{for } M = \{10^3, 10^4, 10^5, 10^6\} \text{ GeV}. \end{aligned}$$

Majorana fermion and B violating operators – but no baryogenesis

Simple way to get n - n bar operator is introduce Majorana fermion X of mass M , coupling to SM by

$$O_6 \sim \frac{1}{\Lambda^2} X u d d$$



Problems creating baryogenesis:

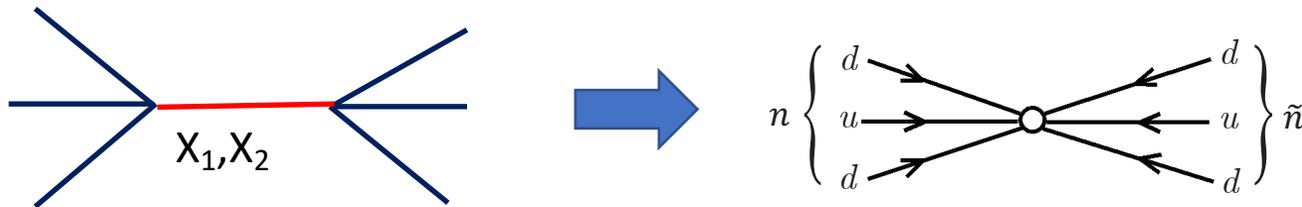
- Nanopoulos-Weinberg theorem: without B-conserving channels no baryon asymmetry
- $2 \rightarrow 2$ process such as $uX \rightarrow \bar{d}\bar{d}$ and $\bar{u}X \rightarrow dd$ have same rate and do not violate CP

Introduce a second Majorana and it can work – “minimal model”

$$\begin{aligned}
 \mathcal{L} \supset & \eta_{X_1} \epsilon^{ijk} (\bar{u}_i^c P_R d_j) (\bar{d}_k^c P_R X_1) \\
 & + \eta_{X_2} \epsilon^{ijk} (\bar{u}_i^c P_R d_j) (\bar{d}_k^c P_R X_2) \\
 & + \eta_c (\bar{u}^i P_L X_1) (\bar{X}_2 P_R u_i) + \text{h.c.}, \\
 & \text{with } |\eta_{X_1}| \equiv \Lambda_{X_1}^{-2}, \quad |\eta_{X_2}| \equiv \Lambda_{X_2}^{-2}, \quad |\eta_c| \equiv \Lambda_c^{-2}.
 \end{aligned}$$

Both X_1 and X_2 mediate n - \bar{n} oscillation — integrating them out at tree level gives

$$c_1 = \frac{1}{(\Lambda_{n\bar{n}}^{(1)})^5} = \frac{1}{M_{X_1} \Lambda_{X_1}^4} + \frac{1}{M_{X_2} \Lambda_{X_2}^4}.$$



Baryogenesis capability

Calculation of the baryon asymmetry — The relevant processes for baryogenesis include

- B violating processes: single annihilation $uX_{1,2} \rightarrow \bar{d}\bar{d}$, $dX_{1,2} \rightarrow \bar{u}\bar{d}$, decay $X_{1,2} \rightarrow udd$, and off-resonance scattering $udd \rightarrow \bar{u}\bar{d}\bar{d}$;
- B conserving processes: scattering $uX_1 \rightarrow uX_2$, co-annihilation $X_1X_2 \rightarrow \bar{u}u$, and decay $X_2 \rightarrow X_1\bar{u}u$;

as well as their inverse and CP conjugate processes. CP

Calculating baryon asymmetry

We solve set of coupled Boltzmann equations for abundances of $X_{1,2}$ and Y_{B-L} above $T=140$ GeV (sphalerons active above 140 GeV) and Y_B below $T=140$ GeV.

Find regions of parameter space where $Y_B=8.6 \times 10^{-11}$

We scan over all the parameters to achieve the proper baryon asymmetry.

Highest priority is getting baryon asymmetry correct – check $n-nbar$ after.

Equal interaction scales gives baryogenesis but not n-nbar signal

$\Lambda_{X_1} \sim \Lambda_{X_2} \sim \Lambda_c$ is “equal interaction scale” case

For $M_{X_1} \sim M_{X_2} > 10^4$ GeV, Λ needs to be high to kick system out of efficient interactions that would otherwise suppress $X_{1,2}$ abundances too much.

→ *Λ too high for n-nbar signal*

For $M_{X_1} \sim M_{X_2} < 10^4$ GeV, Λ still needs to be somewhat high for out-of-equilibrium but then $\varepsilon_{CP} \sim M_{X_2}^2/\Lambda^2$ is too low for baryogenesis.

→ *Cannot work well for baryogenesis when we force down M_{X_i} in this scenario.*

Conclusion: $\Lambda_{X_1} \sim \Lambda_{X_2} \sim \Lambda$ case maximum possible Y_B that also has n-nbar visible at ESS is $O(10^{-13})$, which is two orders of magnitude too low.

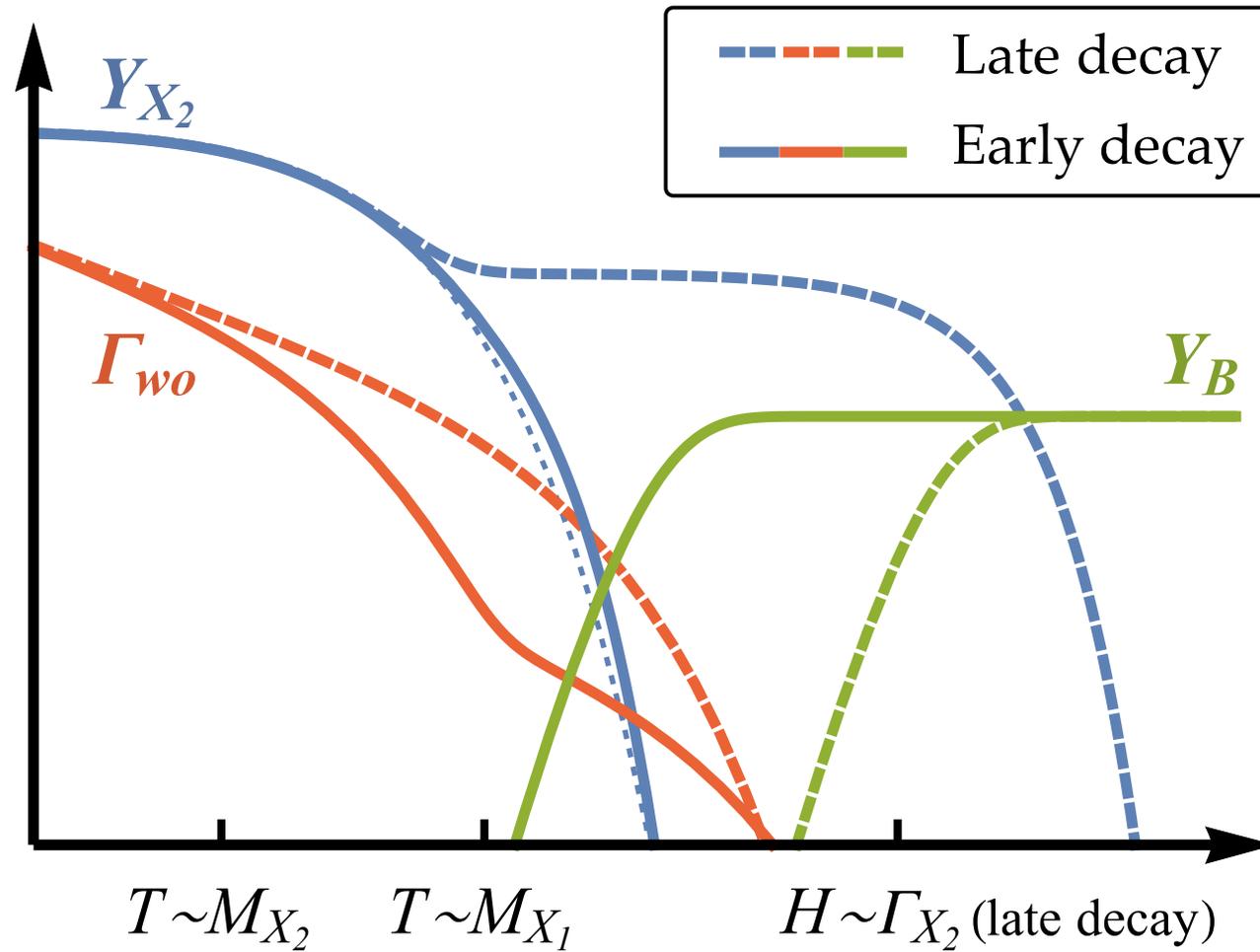
Therefore:

Hierarchy of Λ 's needed for good baryogenesis and visible n-nbar oscillation

- Scenario possible for hierarchy in UV theory
- Or EFT generated at different loop orders

Baryogenesis and n-nbar visibility is compatible in two distinct scenarios:
Late decays of X_2 and earlier decays.

Schematic of Late decay and early decay scenarios for baryon asymmetry



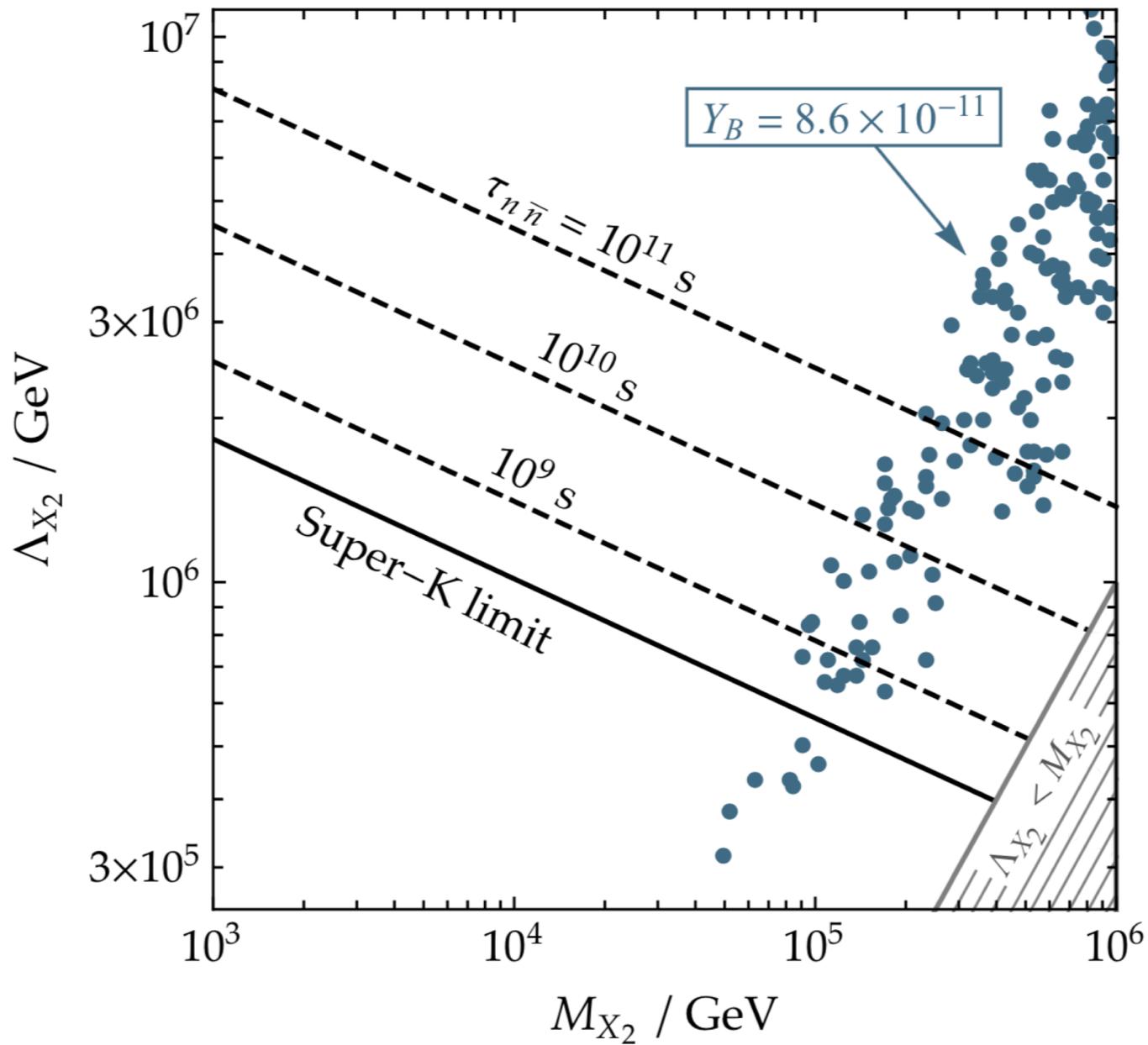


FIG. 3. Parameter space of the minimal EFT probed by $n\text{-}\bar{n}$ oscillation for the early decay scenario, assuming $M_{X_2} = 4M_{X_1}$. Points represent solutions with $Y_B = 8.6 \times 10^{-11}$ found in a scan over $\Lambda_{X_2} < \Lambda_{X_1} < 100\Lambda_{X_2}$, $M_{X_2} < \Lambda_c < \Lambda_{X_2}$. For all these points, $\Lambda_{X_1} \sim 10\Lambda_{X_2}$ is needed to suppress washout. The gray shaded region marks $\Lambda_{X_2} < M_{X_2}$, where EFT validity requires greater than $\mathcal{O}(1)$ coupling.

Grojean, Shakya, JW, Zhang, PRL 2018

Possibility there for discovery of $n\text{-}\bar{n}$ oscillations directly correlated with baryogenesis.

Conclusion

Baryon number conservation is a soft principle that we should expect to be violated.

Baryon violation is needed for baryogenesis – many ideas to implement that.

A minimal, two-state Majorana solution can provide needed baryogenesis

This theory also predicts n - \bar{n} oscillation lifetime

- 1) $\tau < \textit{current limit}$ (ruled out parameter space)
- 2) $\tau > \textit{future projected limits}$ (never will be seen this way – sad)
- 3) $\textit{Current limit} < \tau < \textit{future projected limits}$ (discovery! – how likely?)